Tentamen Voortgezette Mechanica

NS-350B, Blok 2, Final; February 1, 2018

Mark on each sheet clearly your name and collegekaartnummer.

Please use a separate sheet for each problem.

Tip: Read all questions and start with the one you find the easiest. Do not use too much time on any one question!

1 Hamiltonian in a Central Forcefield

A particle moves under the influence of a central force $\vec{F}(\vec{r},t) = -\frac{k}{r^2}e^{-\beta t}\hat{r}$, where k and β are positive and constant, t is time, and r is the distance of the particle to the origin. (total: 15 points)

- a) Using the general definition of a Hamiltonian, find \mathcal{H} of this system. (6 points)
- b) Use the Hamiltonian to find the equation of motion of the particle. (5 points)
- c) Compare the Hamiltonian you calculated to the total energy. Why is it equal/not equal to the total energy T + U? (2 points)
- d) Is the energy of the particle conserved? Use a simple example to discuss why (not). (2 points)

2 Rigid Body Rotations and Chandler Wobble

We are considering rotational motion of a rigid body with homogeneous mass distribution (mass M, density ρ). Let us start with a triangular prism where H < B < L. (total: 22 points)

- (a) Calculate the inertial tensor I of the prism relative to the centre of rotation and the axes shown in the figure! (8 points)
- (b) One of the given axes is a principal axis ("hoofdas").

 Identify this principal axis from the found inertial
 tensor and symmetry considerations, and in doing so define the term "principal axis". Why
 do you find a principle axis for this body though it does not possess rotational symmetry? (4
 points)

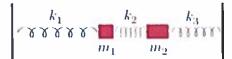
For the second part of this problem we consider a rigid body with only two different principal moments $\lambda_1 = \lambda_2 \neq \lambda_3$. Our *body frame*, as per definition, is fixed in the centre of mass and its axis are pointing along the principal axis of the body.

- (c) Write out the three Euler Equations for this rigid body without external torque. (3 points)
- (d) The rotation of Earth is well described by the Euler Equations that you have found (\hat{e}_3 is the rotational axis of earth, $\lambda_3 \approx 306/305\lambda_1$). The Euler Equations suggest that the direction of the angular velocity $\vec{\omega} = \omega_1 \hat{e}_1 + \omega_2 \hat{e}_2 + \omega_3 \hat{e}_3$ of Earth changes in time. Find the period of this "Chandler Wobble". (4 points)

(e) In reality, the Chandler Wobble has a period of 433 days, which is not the period you have found. Give two potential explanations that could explain this discrepancy - other than that we used the wrong principal axes or principal moments. (3 points)

6cm

Coupled Oscillations 3



- (a) Choose x_1 and x_2 as generalized coordinates for the displacement from the equilibrium positions of the blocks with masses m_1 and m_2 , and find the equation of motion for both blocks. (5 points)
- (b) Combine the set of of motions into matrix form $\mathbf{M}\ddot{\vec{x}} = -\mathbf{K}\vec{x}$. (3 points)

From now on, use $m_1 = 2m$, $m_2 = m$, $k_1 = 4k$, $k_2 = k$ and $k_3 = 2k$.

- (c) Find the frequencies of (small) oscillations of the modes. (6 points)
- (d) Find the normal modes of this system. Describe the physical motions to which the normal modes correspond. (4 points)

4 Multiple Choice

The final questions for this exam are multiple choice (on a separate sheet). Please make sure to remove the staple *cleanly* before you hand in the answer sheet. [total: 5+2 bonus points]

Useful Formulas

$$\sin^2\alpha + \cos^2\alpha = -e^{i\pi} = 1, \qquad T_0 \approx 2\pi\sqrt{\frac{\ell}{g}} \qquad \text{hoop: } I_{zz} = mr^2$$

$$c = \frac{\ell^2}{\gamma\mu} \quad E = \frac{\gamma^2\mu}{2\ell^2} \left(\epsilon^2 - 1\right) \quad \mu = \frac{m_1m_2}{m_1 + m_2} \quad \frac{c_1}{1 + \epsilon_1\cos\left(\phi + \delta_1\right)} = \frac{c_2}{1 + \epsilon_2\cos\left(\phi + \delta_2\right)}$$

$$U_{\text{cf}}(r) = \frac{\ell^2}{2\mu r^2} \quad \mathcal{L}_{\text{rel}} = \frac{1}{2}\mu \ddot{r}^2 - U(r) \qquad J_{ij} = I_{ij}^{\text{cm}} + m\left(||\vec{a}||^2 \delta_{ij} - a_i a_j\right)$$

$$\mathbf{I} = \iiint dV \ \rho(x, y, z) \begin{pmatrix} y^2 + z^2 & -xy & -xz \\ -yx & x^2 + z^2 & -yz \\ -zx & -zy & x^2 + y^2 \end{pmatrix}$$

$$\mathcal{L} = T - V \qquad \frac{\partial \mathcal{L}\left(q(t), \dot{q}(t), t\right)}{\partial q} - \frac{d}{dt} \frac{\partial \mathcal{L}\left(q(t), \dot{q}(t), t\right)}{\partial \dot{q}} = 0$$

$$\mathcal{H} = \sum p_i \dot{q}_i - \mathcal{L} = T + U \qquad \dot{p}_i = -\frac{\partial \mathcal{H}}{\partial q_i} \quad \dot{q}_i = \frac{\partial \mathcal{H}}{\partial p_i}$$

$$\vec{L} = (\lambda_1 \omega_1, \lambda_2 \omega_2, \lambda_3 \omega_3) \quad \dot{\vec{L}} + \vec{\omega} \times \vec{L} = \vec{\Gamma} \qquad \frac{1}{\sqrt{1+x}} = 1 - \frac{x}{2} + \mathcal{O}(x^2)$$

$$\lambda \vec{\omega} = \mathbf{I} \vec{\omega} \qquad (\mathbf{I} - \lambda \mathbf{E}_3) \vec{\omega} = 0 \qquad \det(\mathbf{I} - \lambda \mathbf{E}_3) = 0 \qquad \mathbf{E}_3 = \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$F_i(d\vec{A}) = \sum_i \sigma_{ij} dA_j \qquad \vec{F}(d\vec{A}) = \Sigma d\vec{A} = \Sigma \hat{n} dA$$

For a surface and a point on that surface: $S(x,y,z)=0,\ P\in S: \quad \vec{n}=\nabla S$ $\int_0^{2\pi}\sin^n(x)\cos^m(x)dx=0 \text{ for odd } m \text{ or } n. \qquad \qquad \int_0^{2\pi}\cos^2(x)dx=\int_0^{2\pi}\sin^2(x)\ dx=\pi$

$$\arctan 0 = 0 \qquad \arctan \pm \infty = \pm \frac{1}{2}\pi \qquad \det(\mathbf{K} - \omega^2 \mathbf{M}) = 0$$

$$f(x^*) = x^* \qquad f(x_a) = x_b \wedge f(x_b) = x_a \Rightarrow f(f(x^*)) = x^* \qquad x_t = x^* + \epsilon \Rightarrow \epsilon_{t+1} \approx f'(x^*) \epsilon_t$$

$$f^{(n)}(x) = \underbrace{f(f(\dots f(x) \dots))}_{\text{ntimes}}$$

Voortgezette Mechanica (NS-350B)

Final Exam Multiple-Choice Part

No documents allowed. The use of electronic devices is forbidden.

Questions using the sign \(\bigap \) may have zero, one or several correct answers and are worth 3 points.

All other questions have a single correct answer and are worth 2 points.

Question 1 For small oscillations (linearized case), the motion of a coupled oscillator with n degrees of freedom can be written using $n \times n$ "mass" and "spring-constant" matrices:

$$\mathbf{M}\mathbf{\ddot{\ddot{q}}} = -\mathbf{K}\mathbf{\ddot{q}}.$$

We can find the normal frequencies ω associated with these small oscillations by solving:

Question 2 The Euler Angles θ , ϕ , and ψ allow us to connect the Euler equation in the body frame to an inertial frame. It turns out that θ is moving in an effective potential, which causes a nodding motion called ...

A Chandler wobble. B free rotation. C steady precession. D free precession. E nutation.

Question 3 A 'boost' in special relativity refers to ...

- A ... the additional factor in the relativistic velocity-addition formula applied to the component parallel to the relative motion.
- B ... the additional factor in the relativistic velocity-addition formula applied to the components perpendicular to the relative motion.
- C ... any Lorentz transformation that leaves corresponding axes parallel.
- ① ... the change of length (length contraction) of an object perpendicular to its relative velocity, when measured from an inertial system at rest.
- $\stackrel{\frown}{E}$... the time delay Δt of two events which appear simultaneous in a frame at rest, when measured in a moving frame.

Question 4 We found that transverse motion on an infinite string leads to the wave equation. Which statements of the wave equation are correct?

- A The wave equation has no solutions in more than two dimension.
- B Standing waves do not travel along the string, their spatial and temporal functions are separated.
- C For the sinusoidal f and g we find standing waves as a solution.
- \bigcirc All solutions to the wave equation on an infinite string are of the form f(x+ct)+g(x-ct), travelling left (or right) on the string.
- E The only solution to the wave equation on a *finite* string is a standing wave with one fixed frequency, depending on L.

| Question 5 Starting from different initial condition, we can write the separation of the two resulting solutions as $ \Delta \varphi(t) \sim e^{\lambda t}$. The coefficient λ is known as the | |
|---|---------------------|
| (A) Control parameter | D Lagrange exponent |
| B Feigenbaum constant | E Poincaré exponent |
| C Liapunov exponent | |
| Question 6 & Coupled oscillations can be described in terms of <i>normal modes</i> . Which of these statements concerning normal modes are true? | |
| Any periodic motion of a coupled oscillator can be written as a (weighted) sum of normal modes. | |
| B Every coupled oscillator has exactly one normal mode. | |
| \bigcirc A normal mode is any motion in which all n coordinates oscillate sinusoidally with the same frequency ω . | |
| \bigcirc In a symmetric case (same spring constant and masses), all n normal modes of a coupled oscillator have the same ω . | |
| (E) Any periodic motion of a coupled oscillator has to be exactly one of the normal modes. | |
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| Question 7 ♣ Which of the following statements apply to the Inertia Tensor of an arbitrary rigid body? | |
| (A) The Inertia Tensor is a function of the shape of the body only and does not depend on the choice of origin or axes. | |
| (B) Every Inertia Tensor has three distinct eigenvalues or <i>principal moments</i> which are different from each other. | |
| C An axis of rotational symmetry is always a principal axis of a body. | |
| D Every Inertia Tensor can be diagonalized. | |
| \bigcirc All elements of an Inertia Tensor I_{ij} have to be positive. | |
| Question 8 A question about the order of magnitude of time dilation. Suppose a certain event happens at intervals of exactly one hour in an inertial frame moving at a steady speed of $300\mathrm{ms^{-1}}$. When measured from an inertial frame at rest, the interval will appear | |
| Alonger by 2 nanoseconds. | |
| B exactly the same length in time, as both are inertial frames. | |
| C shorter by half a nanosecond. | |
| Dlonger by half a millisecond. | |
| Eshorter by 2 milliseconds. | |
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