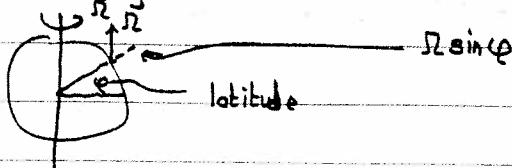


answers 6FD exam d.d. 10-11-2010

- 2a. Coriolis parameter / planetary vorticity : $f = 2\Omega \sin \varphi$ Ω angular speed of rot.
 $2 \times \frac{1}{2}$ angular speed of rotation of fluid particle around local z-axis:



1b. Richardson number $R_i = \frac{\Delta \rho g H}{\frac{1}{2} \rho_0 U^2}$

where $\Delta \rho$ density difference between top-bottom, H mean thickness,
 g acceleration due to gravity, ρ_0 mean density, U velocity scale.

Meaning: ratio of potential energy (due to stratification) and kinetic energy.

In this case: $\Delta \rho = 0$ (equation is that of homogeneous shallow water model)
 So $R_i \approx 0$. So indeed, very small.

c. Pressure gradient force per mass unit in the y-direction.

To be derived from $-\frac{1}{\rho} \frac{\partial p}{\partial x}$ (see equation sheet).

Shallow water \rightarrow hydrostatic balance, so

$$\frac{\partial p}{\partial z} = -\rho_0 g \Rightarrow p = \rho_0 g (\eta - z) + p_0$$

using boundary condition $p = p_0$ at $z = \eta$.

Assuming $p_0 = \text{constant}$: $\frac{\partial p}{\partial x} = \rho_0 g \frac{\partial \eta}{\partial x}$, so $-\frac{1}{\rho_0} \frac{\partial p}{\partial x} = -g \frac{\partial \eta}{\partial x}$

Remark: you can't say $p = \rho_0 g \eta$ This only applies at level $z = 0$. Here use of
hydrostatic balance is essential!

d. Homogeneous shallow water model: u, v independent of z ,
 w varies linearly with z .

So $v = 1 \text{ ms}^{-1}$ at $z = b$ and $w = u \frac{\partial b}{\partial x} + v \frac{\partial b}{\partial y}$ at $z = b$
 or $w = v \frac{\partial b}{\partial y} = 0.02 \text{ ms}^{-1}$

e. $v = 1 \text{ ms}^{-1}$ at $z = \eta$ (depth uniform).

$$w = w_{\text{bot}} + (w_{\text{surf}} - w_{\text{bot}}) \frac{(z-b)}{h}$$

We know $w_{\frac{1}{2}} = w_{\text{bot}} + \frac{1}{2}(w_{\text{surf}} - w_{\text{bot}}) = \frac{1}{2}w_{\text{bot}} + \frac{1}{2}w_{\text{surf}}$

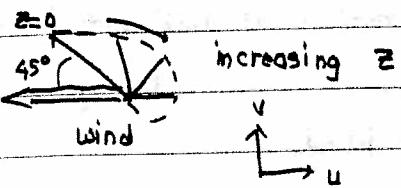
So $w_{\text{surf}} = 2w_{\frac{1}{2}} - w_{\text{bot}} = 0.08 \text{ ms}^{-1}$

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2a. At $y=0$ $\tau^x = -T_b + \hat{T} < 0$

No geostrophic flow in interior; Northern Hemisphere

So



note: flow speed decays exponentially
on scale $d = \left(\frac{2v_E}{f}\right)^{1/2}$: Ekman depth

b. The equations are

$$-fv = v_E \frac{\partial u}{\partial z} \quad \text{bound. conditions } v_E \frac{\partial u}{\partial z} = \frac{T^x}{\rho_0}, v_E \frac{\partial v}{\partial z} = \frac{T^y}{\rho_0} \text{ at } z=0$$

$$fu = v_E \frac{\partial v}{\partial z}$$

$u, v \rightarrow 0$ for $z \rightarrow -\infty$

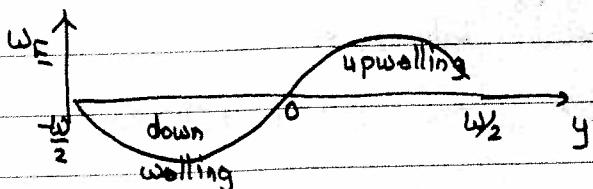
Note: no pressure gradient force here (no geostrophic flow)

Integrate and use $\bar{U} = \int u dz$ at $z=0$ and $\bar{V} = \int v dz$

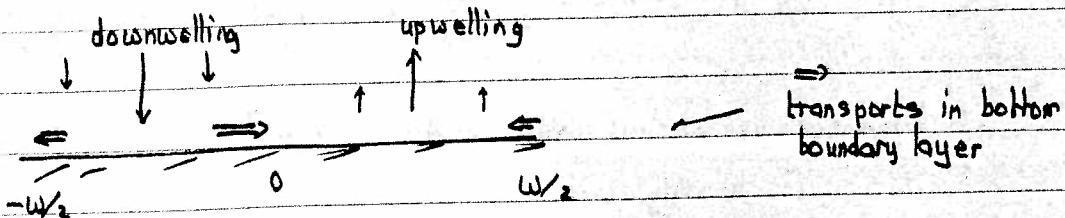
$$-fT = v_E \frac{\partial u}{\partial z} \Big|_{-\infty} = \frac{T^x}{\rho_0 f} \quad \text{so} \quad \bar{v} = \frac{T_b - \hat{T}}{\rho_0 f}$$

Likewise, $\bar{U} = 0$

c. $w_E = -\frac{1}{\rho_0 f} \frac{\partial T^x}{\partial y} = \frac{+T}{\rho_0 f} \frac{2\pi}{\omega} \sin\left(\frac{2\pi y}{\omega}\right)$



d.



So high pressure in southern region, low pressure in north

e. From eq. sheet: $\frac{d}{2} \bar{\xi} = +\frac{1}{\rho_0 f} \left(\frac{\partial T^y}{\partial x} - \frac{\partial T^x}{\partial y} \right)$

Here $\frac{-d}{2} \frac{\partial u_g}{\partial y} = -\frac{1}{\rho_0 f} \frac{\partial T^x}{\partial y}$ T^x and apply $u_g > 0$

So $u_g = \frac{2 T^x}{\rho_0 f d} + \text{constant} \Rightarrow u_g = \frac{2}{\rho_0 f d} \hat{T} \cos\left(\frac{2\pi y}{\omega}\right)$
 \Rightarrow westward flow near $y = \pm \frac{1}{2} \omega$, but eastward (opposing wind stress!) in central region

f. $\Gamma = \oint \tilde{u}_g \cdot d\tilde{r} \rightarrow \Gamma = u_g(0)L - u_g\left(\frac{\omega}{2}\right)L \leftarrow \dots = \frac{4\hat{T}L}{\rho_0 f d}$