

1 a. Critical level  $z_c$  defined as

$$U_0(z=z_c) = c$$

where  $c = \frac{\sigma}{k}$  is the phase speed of the waves.

In the present case:

$$c_1 = \left( \frac{g}{k_1} + \frac{\Gamma}{e} k_1 \right)^{1/2} \quad \text{and} \quad c_2 = \left( \frac{g}{k_2} + \frac{\Gamma}{e} k_2 \right)^{1/2}$$

So

$$\frac{u_x}{k} \ln \left( 1 + \frac{z_{c,i}}{z_0} \right) = \frac{g}{k_i} + \frac{\Gamma}{e} k_i \quad \text{where } i=1,2$$

$$\Rightarrow \boxed{z_{c,i} = z_0 \left[ -1 + \exp \left\{ \frac{k}{u_x} \left( \frac{g}{k_i} + \frac{\Gamma}{e} k_i \right)^{1/2} \right\} \right]}$$

b. Miles shear instability mechanism.

Positive feedback between ocean and atmosphere.

Waves at sea surface create perturbations in flow in atmospheric boundary layer



These perturbations lead to transfer of horizontal momentum in vertical direction.

Divergence of this momentum transport only at critical level  $z=z_c$ .

So here mean wind is affected; loss of momentum and loss of energy density of atmospheric flow.

This loss of energy is a gain of energy density of the water:

wave amplitude increases exponentially

c. The unit of  $F$  is  $m^2/Hz$  or  $m^2 s$ :  $F$  is the wave variance per frequency.

In other words  $\iint F d\omega d\theta = \langle \sigma^2 \rangle$  where  $\langle \sigma^2 \rangle$  is wave variance.

Unit of  $\sigma$  is  $s^{-1}$ , so unit of  $\beta$  is  $m^2 s^{-4}$

d. Energy density  $\langle E \rangle = \rho g \langle \sigma^2 \rangle$  for gravity waves

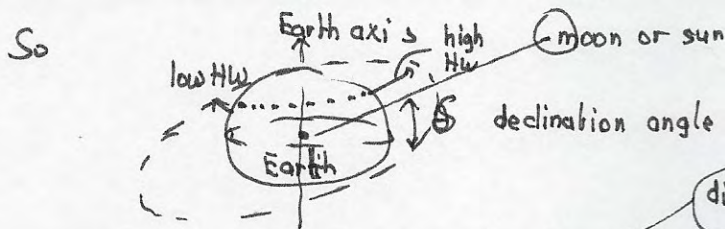
$(\rho g + \Gamma k^2) \langle \sigma^2 \rangle$  for capillary gravity waves

(but capillary waves have hardly energy at this stage)

$\Rightarrow$  Energy density of high frequency waves

$$\langle E \rangle_{\text{high}} = \rho g \int_{\frac{\omega}{\omega_0} - \frac{\pi}{2}}^{\frac{\omega}{\omega_0} + \frac{\pi}{2}} \int_0^{2\pi} \beta \sigma^{-5} \cos^2(\theta - \theta_\omega) d\theta d\sigma = \rho g \frac{1}{2} \pi \frac{1}{4} \sigma_p^{-4}$$

- 2a. Dominant period: 1 HW, 1 LW per day, so period  $\sim 24$  h: diurnal tide.  
 Diurnal tides result from declination: line connecting centre of mass Earth and that of celestial body is not in equatorial plane.



The equilibrium tide (resulting from tidal force balancing pressure gradient force is indicated by the dashed curve. difference local gravitational force and gravitational force in centre of mass Earth

Thus an observer at latitude  $\varphi$  on rotating Earth will experience in 1 cycle of  $\sim 24$  h a high HW and low HW. Thus semi-diurnal + diurnal tide.

The smaller  $\delta$ , the weaker the diurnal tide.

- b. Pure coincidence, the magnitude of the diurnal tide is not determined by full moon, last quarter etc, but by the declination angle  $\delta$ . Apparently (and indeed) on 27/10/15 the declination angle of the moon was small. So was that of the sun (the latter was 0 on 21/09/15 and will become  $23.5^\circ$  on 21/12/15)

During full moon: period  $\sim 12^{\text{h } 25^{\text{m}}}$ , the semi-diurnal tides are <sup>close to</sup> maximum at that time.

- c. Yes, in particular because the distance moon - Earth will be different. That distance varies on a period of 27.5 day and it causes tidal forces to be stronger when the distance is small, and weaker when the distance is large.

From moon data: distance larger in that period  $\rightarrow$  tidal range smaller



3a. Two types of waves:

Kelvin waves:  $\sigma^2 = gHk^2$  so  $\lambda = \frac{2\pi}{k} = \sqrt{gH} T$

where  $T = 2\pi/\sigma$  the period

Poincaré waves:  $\sigma^2 = f^2 + gH(k^2 + \frac{n^2\pi^2}{B^2})$

so  $\lambda = \frac{2\pi}{k}$  and  $k_n = \left( \frac{\sigma^2 - f^2}{gH} - \frac{n^2\pi^2}{B^2} \right)^{1/2}$

$T = 44700 \text{ s}, \sigma = 1.4 \cdot 10^{-4} \text{ s}^{-1}$

So  $\lambda_{\text{Kelvin}} \approx 7070 \text{ km}$

Furthermore  $k_1 = (3.84 \cdot 10^{-13} - 4 \cdot 10^{-13})^{1/2} \text{ m}^{-1} \ll \text{no solution}$   
So no travelling Poincaré waves

However, when  $\left( \begin{matrix} g = 9.81 \text{ m s}^{-2} \\ \sigma = 1.405 \cdot 10^{-4} \text{ s}^{-1} \end{matrix} \right)$  is used:  $n=1$  Poincaré wave is present

(with an extremely large wavelength)

- b.
- Kelvin waves, trapped on east/west coast, length scale  $R = \frac{\sqrt{gH}}{f}$
  - Poincaré waves, trapped on north coast

Rossby radius of deformation  
 $\sim 1600 \text{ km}$

length scale  $l_n = s_n^{-1}$   
from dispersion relation

$$\sigma^2 = f^2 + gH \left( -s_n^2 + \frac{n^2\pi^2}{B^2} \right) \rightarrow s_n = \left( \frac{f^2 - \sigma^2}{gH} + \frac{n^2\pi^2}{B^2} \right)^{1/2}$$

Note  $s_1 < s_2 < \dots$

Here  $s_1 \approx (1.6 \cdot 10^{-14})^{1/2} \text{ m}^{-1} \approx 1.26 \cdot 10^{-7} \text{ m}^{-1}$  if  $\sigma = 1.4 \cdot 10^{-4} \text{ s}^{-1}$  is used

So  $l_1 \approx 8000 \text{ km}$  (so they occur in quite a large part of the domain)

If  $\left( \begin{matrix} g = 9.81 \text{ m s}^{-2} \\ \sigma = 1.41 \cdot 10^{-4} \text{ s}^{-1} \end{matrix} \right)$  is used:  $n=1$  is not trapped

$\Rightarrow$  compute  $s_2 \approx 1.1 \cdot 10^{-6} \text{ m}^{-1} \rightarrow l_2 = s_2^{-1} \approx 900 \text{ km}$